

# **Low cost robotic imaging system for high precision photometry**

**Olivier Guyon (Subaru Telescope & Univ. of Arizona)  
Josh Walawender (Univ. of Hawaii, VYSOS project)**

Project made possible thanks to :

Bo Reipurth (Univ. of Hawaii, VYSOS project)  
Mauna Loa Observatory

With help of:

Paul Stewart (Univ. of Sydney)

# System characteristics / summary

Experimental system, main goals are:

- test cost-effective approach to surveys (unit cost = \$14000 with labor)
- test suitability of mass produced CMOS arrays for science
- test new efficient algorithm for exoplanet transit detections

150 sq degree FOV, 7cm aperture (etendue = 1m telescope, 1deg diam FOV)

10" per pixel

photon-noise limited on sky background

**Low cost, use of mass-produced commercial components:**

- Canon DSLR camera
- Atlas EQ-G mount
- 85mm F1.2 lens

No dome, No custom electronics or machining

Open source software, running under conventional linux

All information available on public website, to encourage duplication & improvements:

[www.naoj.hawaii.edu/staff/guyon](http://www.naoj.hawaii.edu/staff/guyon)

Fully robotic, robust to weather, hardware failures, software errors

automatic decision making (flats, darks, observe, choose target)

**Easy to duplicate and upgrade, low cost – scalable to multiple units, higher angular resolution (with longer focal length lens)**

Installed Dec 30, 2010 – working robotically now

New algorithms for high precision photometry (looking for exoplanet transit)

# Key components



\$1920



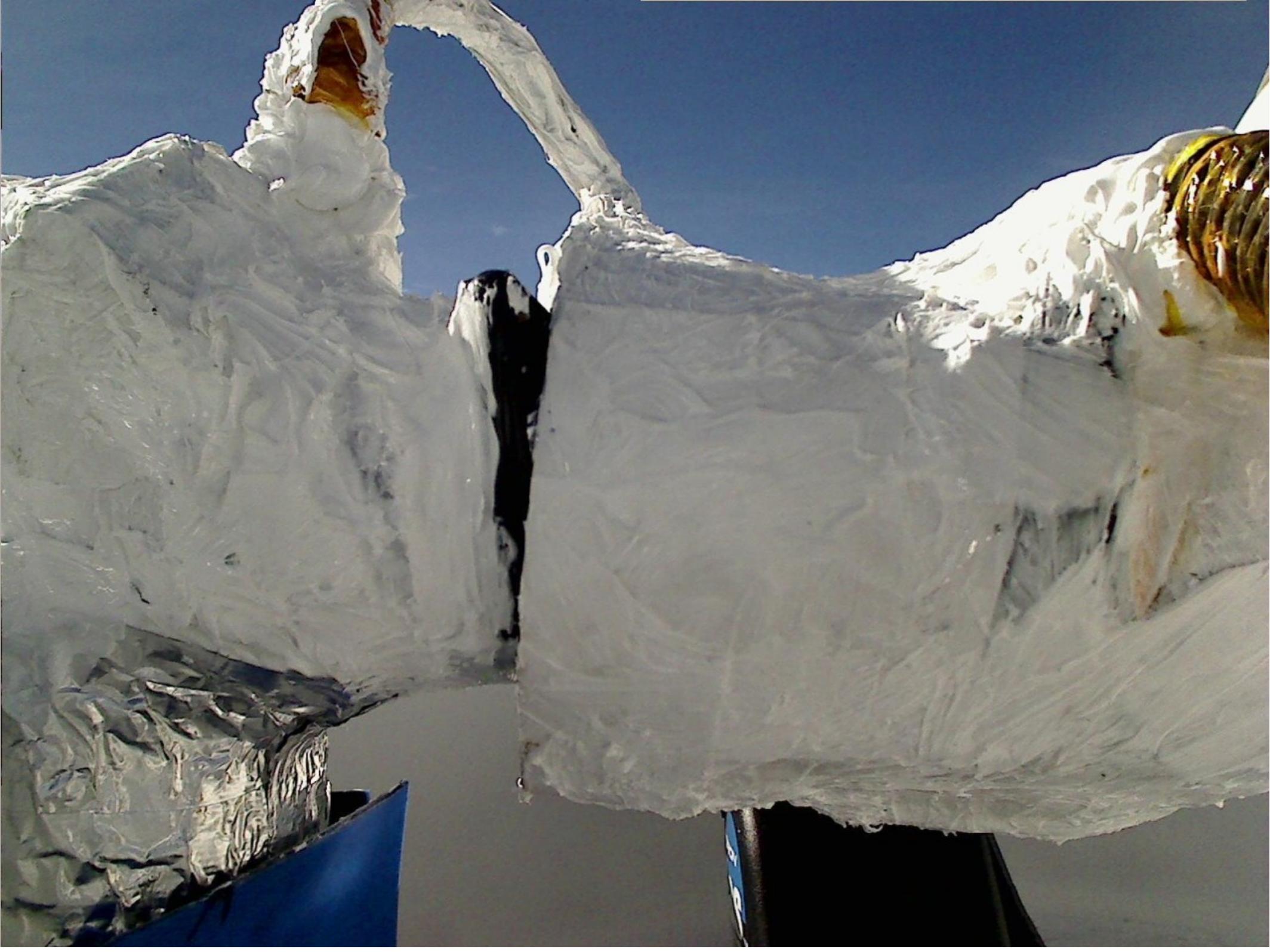
\$550



\$1400

Laptop  
\$800





# **New high precision photometric algorithm for transiting exoplanets**

Exoplanet transit hypothesis tested for each transit period/phase/duration

→ more optimal than conventional process: light curve → transit identification

For each test of transit, use data around transit time to build optimal linear combination of field stars which reproduces target star image (spatial and temporal)

- optimally takes into account color effects, subpixel effects
- good correction for variable extinction

Overcomes disadvantages of color array, and uses color information to reach higher precision

*Conventional process (light curve → transit identification) does not work well because the algorithm used to compute light curve should be a function of the transit parameters*

**This approach is required for transit detection with color array, and will also benefit projects using conventional B/W CCDs**